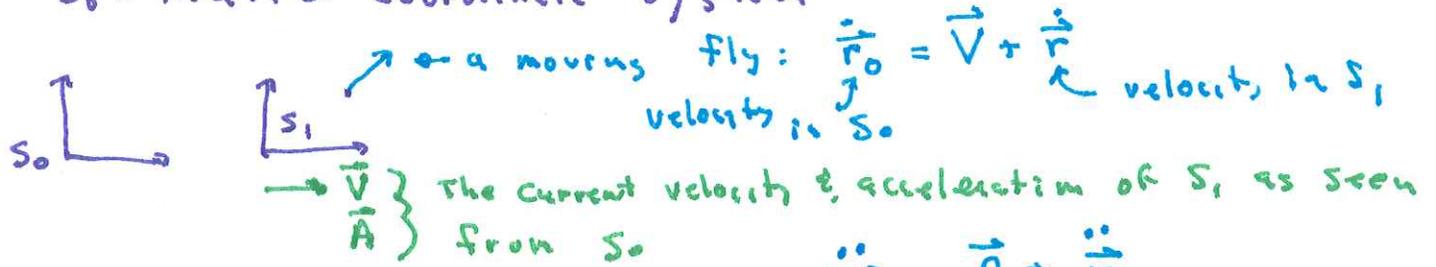


View in a non-inertial coordinate system ($= S_1$)

$S_0 =$ inertial coordinate system



$$\dot{\vec{r}}_0 = \vec{V} + \dot{\vec{r}}_1$$

velocity in S_0

$$\ddot{\vec{r}}_0 = \vec{A} + \ddot{\vec{r}}_1$$

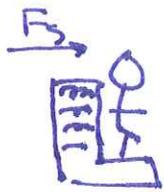
acceleration in S_0

Now: $m \ddot{\vec{r}}_0 = \vec{F}$

$\left. \begin{matrix} \vec{A} + \ddot{\vec{r}}_1 \end{matrix} \right\}$

$$m \ddot{\vec{r}}_1 = \vec{F} - m \vec{A}$$

a pseudo force "inertial force" that like gravity is proportional to m



S_1 view

Force balance: inertial force back, Spring force forward No motion in my frame.

So view: the springs in the car seat are compressed because they are supplying the force that accelerates me

Special Case: uniformly rotating frame - Ω $\frac{\text{rads}}{\text{sec}}$

Remark: $\omega = \Omega$ denote the rotational velocity - $\frac{\text{rad}}{\text{sec}}$ or $\frac{\text{deg}}{\text{sec}}$ or rpm

The magnitude of the vector is that rate

The direction is given by right hand rule.

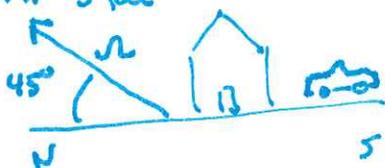
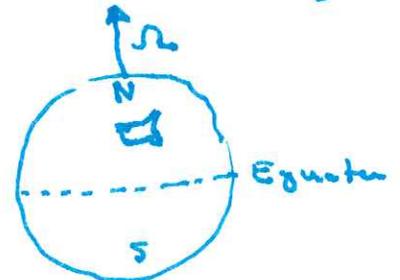
Eg



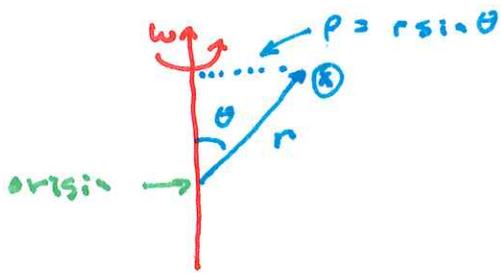
For this spinning disk $\vec{\omega}$ is out-of-page

Often the rotating frame is Earth

Here in MN $\vec{\omega}$ points toward North Star



Consider a point P & a vector \vec{r} pointing from origin (on spin axis) to that point:



$$\text{Speed of point} = \frac{2\pi P}{T} = \omega P = \omega r \sin \theta$$

Period T

direction of velocity = into page

$$\vec{v} = \vec{\omega} \times \vec{r} \text{ gives correct magnitude/direction}$$

The above derivation applies to any rotating vector \vec{Q}

$$\frac{d}{dt} \vec{Q} = \vec{\omega} \times \vec{Q}$$

Questions: what if in the rotating frame S_1 , \vec{Q} shows change?

Let \hat{e}_i be the coordinate axes in the rotating frame S_1

Note: $\frac{d}{dt} \hat{e}_i = \vec{\Omega} \times \hat{e}_i$

↖ assumed constant rotation speed of frame S_1

↖ time rate of change of \hat{e}_i in Frame S_0

$$\vec{Q} = \sum Q_i \hat{e}_i$$

↖ the components of \vec{Q} in the rotating frame (I'm assuming these are changing)

$$\left(\frac{d\vec{Q}}{dt} \right)_0 = \sum \frac{dQ_i}{dt} \hat{e}_i + \sum Q_i \vec{\Omega} \times \hat{e}_i = \left(\frac{d\vec{Q}}{dt} \right)_1 + \vec{\Omega} \times \vec{Q}$$

↑ in S_0

↑ in S_1

Apply twice to \vec{r} to get acceleration:

$$\left(\frac{d}{dt} \vec{v} \right)_0 = \left(\frac{d}{dt} \vec{v} \right)_1 + \vec{\Omega} \times \vec{v} = \left(\frac{d^2 \vec{r}}{dt^2} \right)_1 + \vec{\Omega} \times \left(\frac{d\vec{r}}{dt} \right)_1 + \vec{\Omega} \times \left(\frac{d\vec{r}}{dt} \right)_1 + \vec{\Omega} \times (\vec{\Omega} \times \vec{r})$$

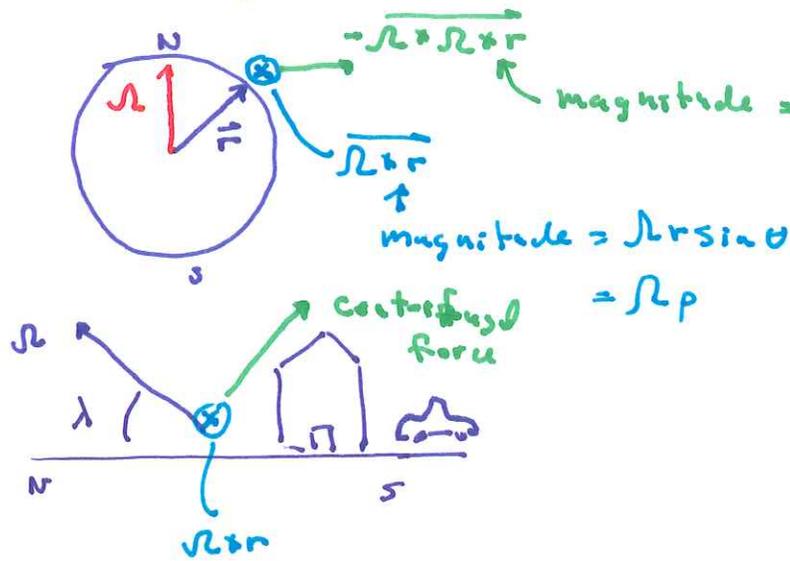
$$\left(\frac{d}{dt} \vec{r} \right)_0 = \left(\frac{d\vec{r}}{dt} \right)_1 + \vec{\Omega} \times \vec{r}$$

↖ in S_1

$$= \left(\frac{d^2 \vec{r}}{dt^2} \right)_1 + 2 \vec{\Omega} \times \vec{v} + \vec{\Omega} \times \vec{\Omega} \times \vec{r}$$

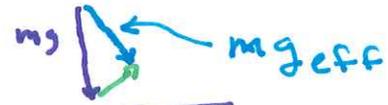
$$m \left(\frac{d^2 \vec{r}}{dt^2} \right)_1 = \vec{F} - m \left(\underbrace{2 \vec{\Omega} \times \vec{v}}_{\text{Coriolis force}} + \underbrace{\vec{\Omega} \times \vec{\Omega} \times \vec{r}}_{\text{centrifugal force}} \right)$$

Centrifugal Force $-m \overline{\Omega \times \Omega \times r}$



} cylindrically outward

Note: The centrifugal force like \vec{g} is proportional to m - what we call \vec{g} is really the combo of gravity & centrifugal



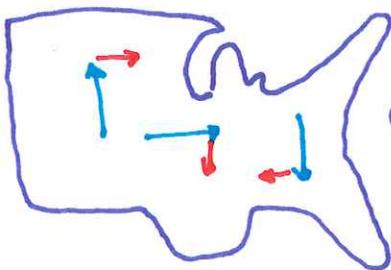
On the equator the centrifugal force is straight up; at the north pole it is zero.

we declare things "level" if they are \perp to \vec{g}_{eff} i.e. balls don't roll off as no net "force"

Coriolis Force. $-2m \overline{\Omega \times V}$ ← like a magnetic field $\vec{B} = -2m\vec{\Omega}$

The vertical component of Ω is $\Omega \sin \lambda$
↑ latitude

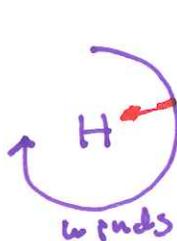
For the following, we ignore the horizontal component of Ω - in the end it does not do much.



⊙ Ω assumed straight up -
 velocity vectors
 Coriolis force vectors

Coriolis force deflects speeding bullets to right

HIGH Pressure systems in US

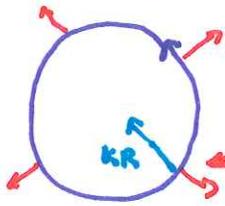


pressure force balanced by Coriolis (in fact a bit less than Coriolis to supply centripetal force for circular motion)

Remark: its an urban legend that Coriolis controls water flow down drains - needs a large scale to have noticeable effect

Foucault Pendulum I - consider conical motion
 (For notational ease spring force $-k\vec{r}$
 replaces gravity $mg\sin\theta$)

$\Omega \odot$

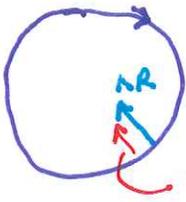


$2m\Omega v$ outward
 ωR where $\omega =$ angular velocity, of conic pendulum

$$KR - 2m\Omega\omega R = m\omega^2 R$$

$$\frac{k}{m} = \omega_0^2 = \omega^2 + 2\omega\Omega \rightarrow \omega_1 = \omega_0 - \Omega$$

approx



$$KR + 2m\Omega\omega R = m\omega^2 R$$

$$\frac{k}{m} = \omega_0^2 = \omega^2 - 2\omega\Omega \rightarrow \omega_2 = \omega_0 + \Omega$$

approx

Note: $\omega_0 \rightarrow$ period of a few seconds
 $\Omega \rightarrow$ period of a day } $\Omega \ll \omega_0$

Given the very small difference in periods - how detect?

Note the system is linear so superposition applies -

Consider superposition of oppositely rotating solutions -

$$e^{i\omega_1 t} + e^{-i\omega_2 t} = e^{i(\frac{\omega_1 - \omega_2}{2})t} \left[e^{i(\frac{\omega_1 + \omega_2}{2})t} + e^{-i(\frac{\omega_1 + \omega_2}{2})t} \right]$$

$$= e^{i(\frac{\omega_1 - \omega_2}{2})t} \left[\cos(\omega_0 t) z \right]$$

$$e^{-i\Omega t} z \cos(\omega_0 t)$$

\uparrow slow rotation
 \uparrow normal back & forth

